

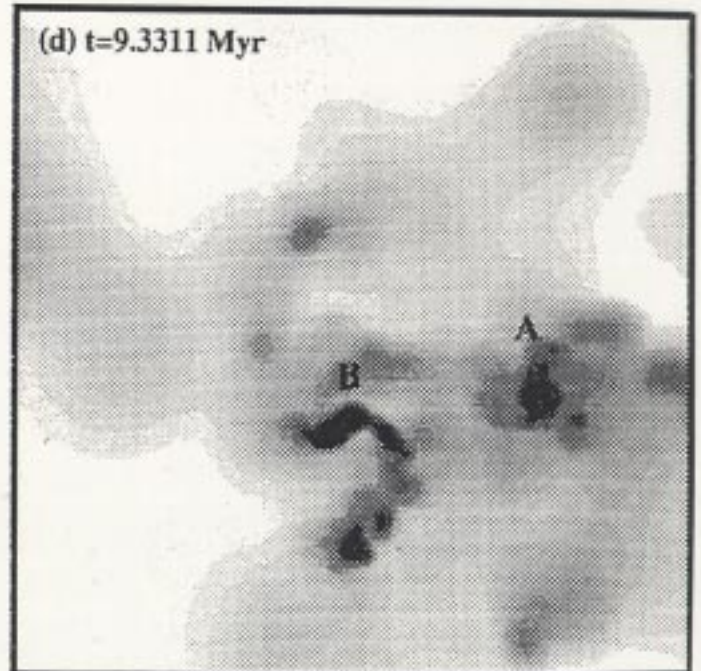
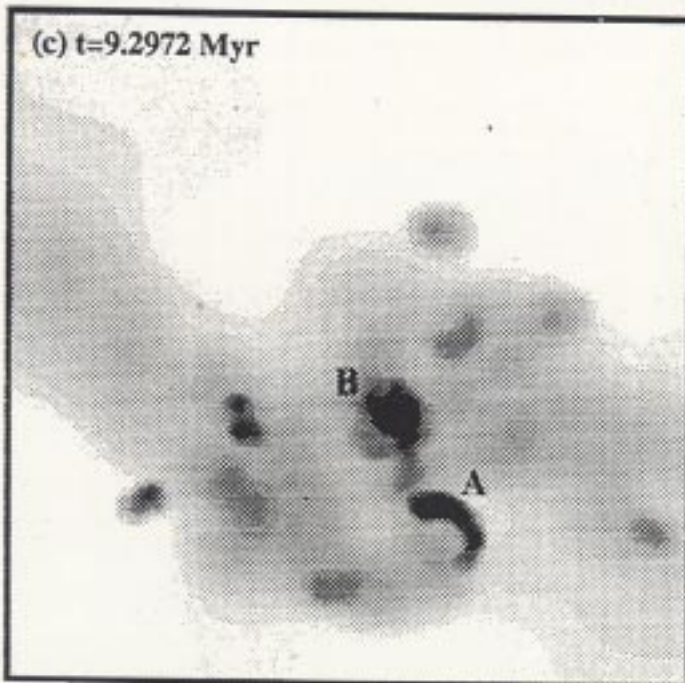
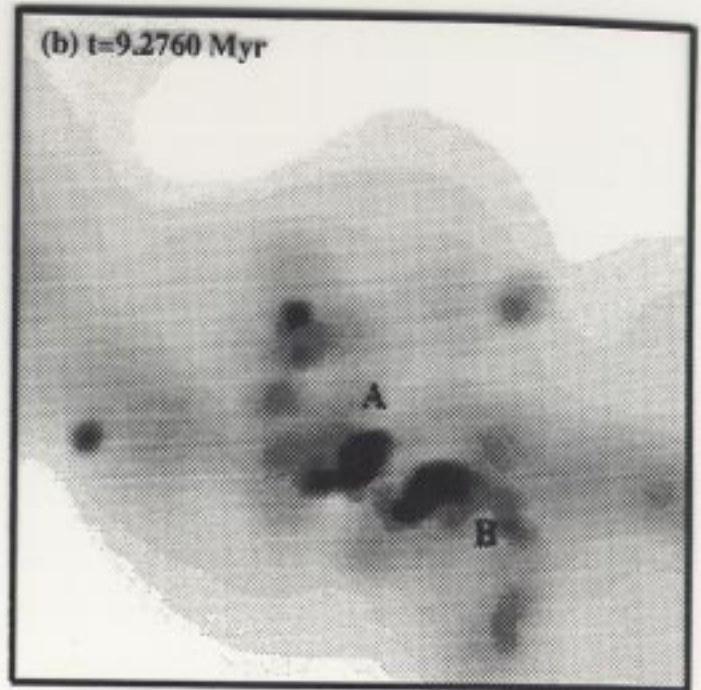
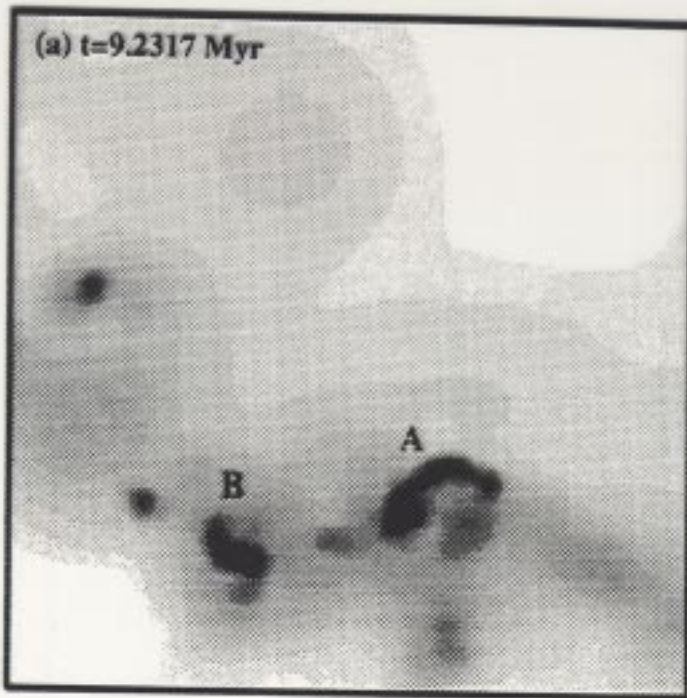
FORMATION MECHANISMS – CLUSTERS

- **Required:** a region that contains many thermal Jeans masses.
- **Problem:** molecular cloud regions on moderate scales (1 parsec) appear to be supported against collapse by magnetic turbulence
- 1. The density is high enough so that ambipolar diffusion time of gas with respect to magnetic field is relatively short (10^6 yr), so magnetic effects could become dynamically unimportant in the whole region. But the required density is 4×10^{-19} g cm⁻³.
- Increase in density could be caused by passage of a shock wave through a cloud, or collision of two clouds followed by cooling in the dense layer behind the shock. Or sudden removal of ionizing flux and cosmic rays from a region causes enhanced magnetic diffusion.
- 2. Magnetic turbulence in molecular clouds decays on a time scale less than the free-fall time according to detailed numerical simulations by Mac Low et al. (1998), Padoan (1998), Stone, Ostriker, and Gammie (1998), Mac Low (1999) unless there is external energy input.
- 3. Turbulence does not decay, but self-gravitating “kernals” can form on a small scale which are cut off from MHD waves in the core (Myers 1998).

Result of collision of 2 clouds, 750 Mo each

each

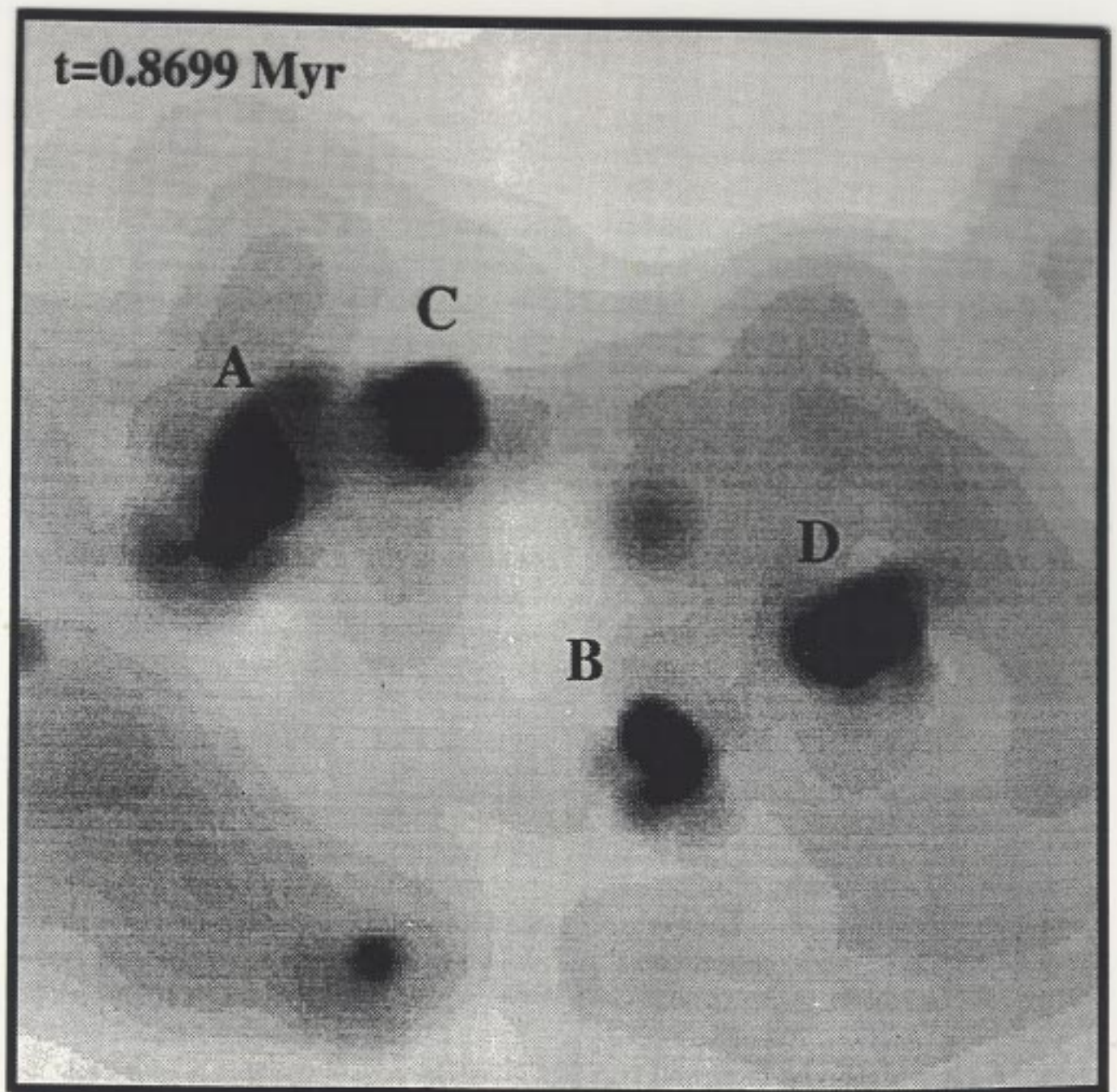
Whitworth et al. M N 277, 727 (1995)



SCALE
0.2 pc

Result of collision of 2 clouds, 75 M_⊙ each

Whitworth et al. MN 277, 727 (1995)



SCALE : 0.14 pc

Klessen and Burkert

CALCULATIONS OF CLUSTER FORMATION

- **Assumptions:** isothermal cloud with gas pressure and gravity
- **Initial conditions:** uniform-density cloud with 222 Jeans masses. Gaussian random density fluctuations are superimposed with power spectrum $P(k) \propto 1/k^2$
- **Numerical technique :** SPH, 500,000 particles
- **Smoothing length** variable in space and time such that the number of “nearest neighbors” for each particle is about 50
- **GRAPE** hardware for solving Poisson equation
- **Periodic boundary conditions**
- **When a highly condensed core forms** it is replaced by a single “sink” particle which continues to accrete gas
- **Results:** formation of knots and filaments, 61 molecular cloud cores appear, with a mass spectrum peaking at the overall Jeans mass for the system

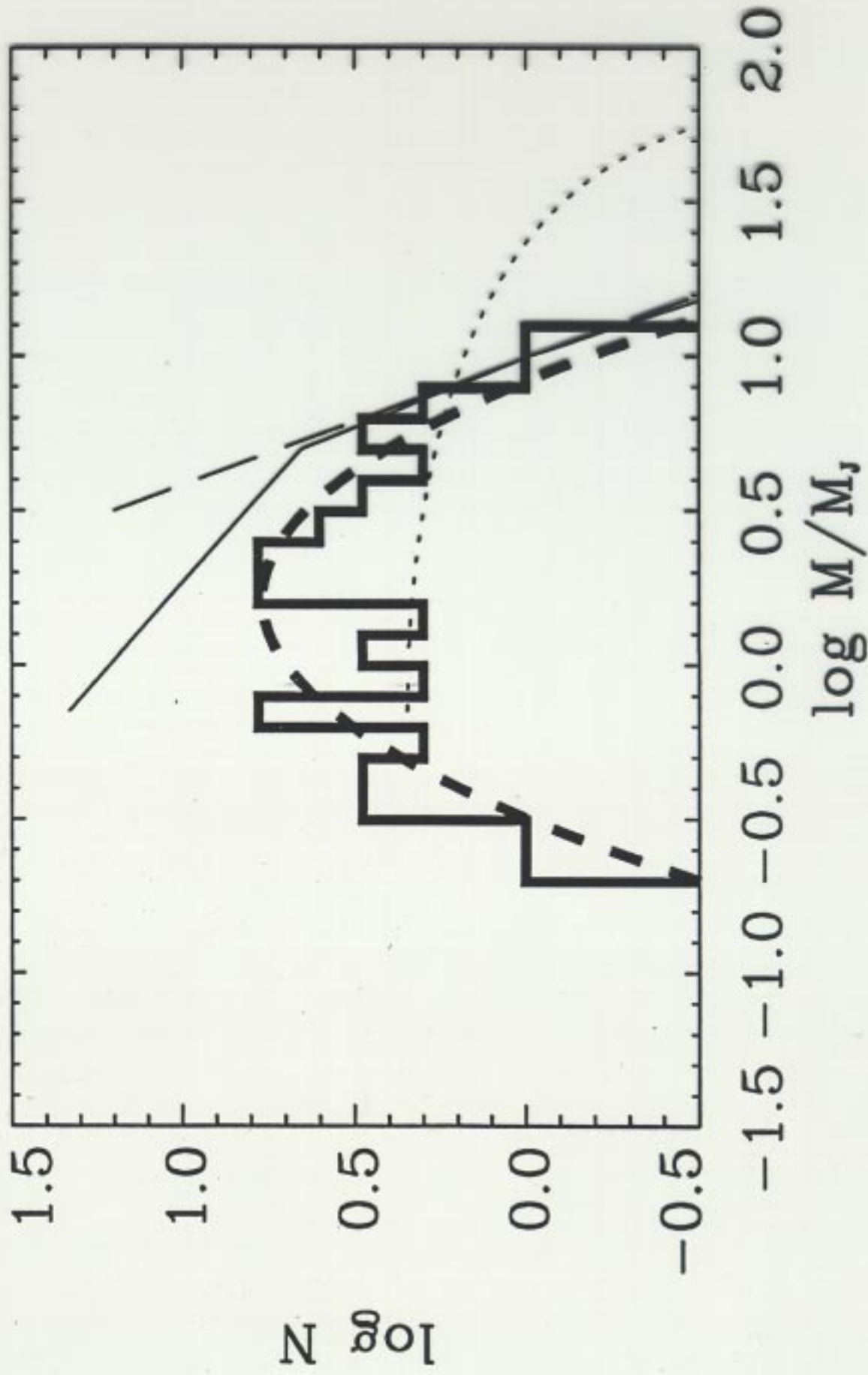
Klessen and Burkert

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Klessen and Burkert 1998

Distribution of core masses



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