



cherenkov  
telescope  
array

**ICRC2017**  
35<sup>th</sup> International Cosmic Ray Conference

# Cherenkov Telescope Array: The Next Generation Gamma-ray Observatory

ICRC 2017 (Busan, Korea), 13 July 2017

The CTA Consortium<sup>1</sup>,  
represented by Rene A. Ong<sup>2</sup>

<sup>1</sup>See [http://www.cta-observatory.org/consortium\\_authors/authors\\_2017\\_07.html](http://www.cta-observatory.org/consortium_authors/authors_2017_07.html)

<sup>2</sup>University of California, Los Angeles, CA, 90095, USA



# 2005-2017: VHE Astronomy Comes of Age

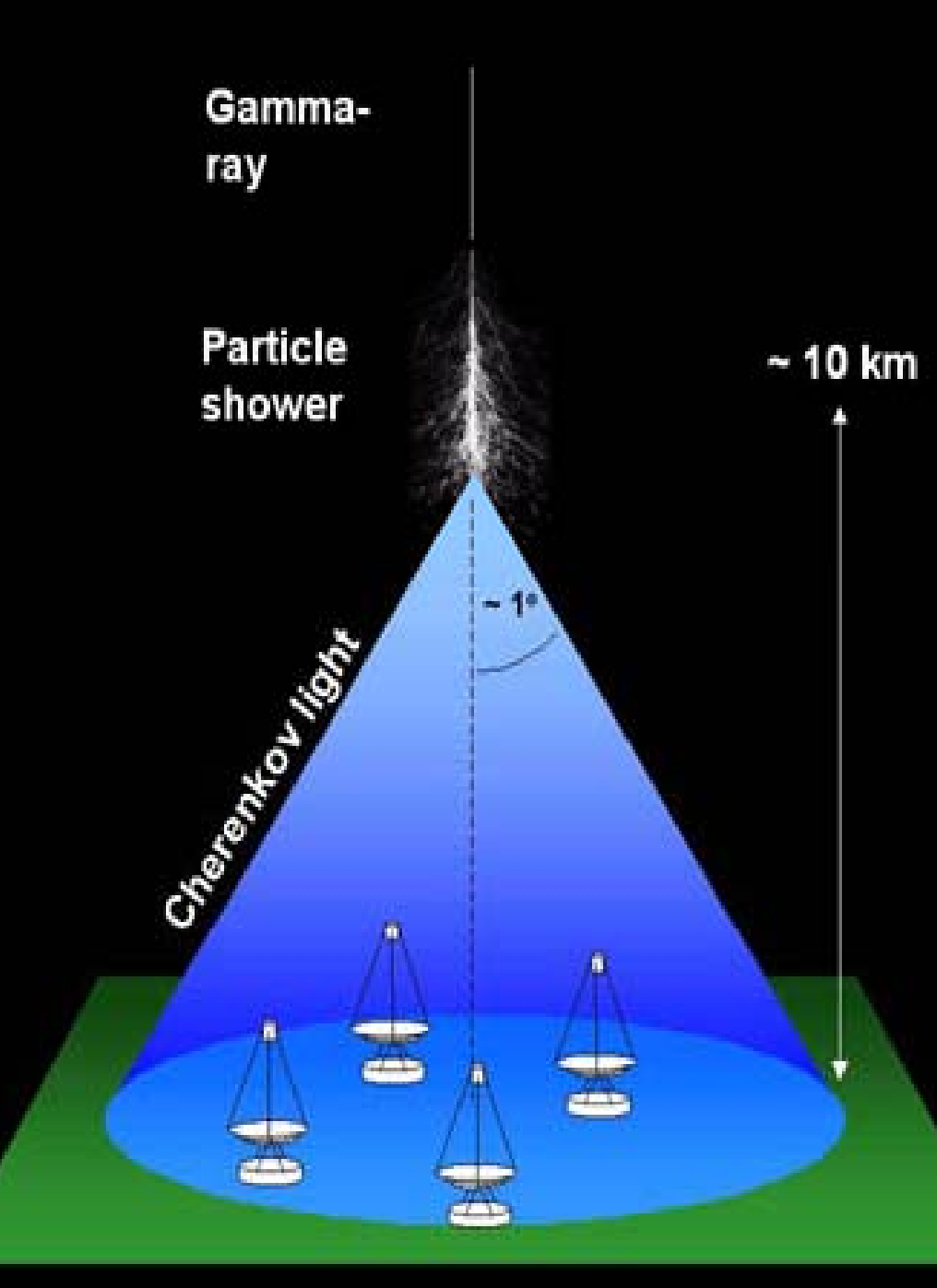
- Dominant expectation (pre-1990)
  - Will find the “cosmic ray” accelerators – probably SNRs
- Reality (2017)
  - Astonishing variety of VHE † emitters
    - Within the Milky Way
      - Supernova remnants
      - Bombarded molecular clouds
      - Stellar binaries - colliding wind & X-ray
      - Massive stellar clusters
      - Pulsars and pulsar wind nebulae
      - Supermassive black hole Sgr A\*
      - Diffuse & extended emission
    - Extragalactic
      - Starburst galaxies
      - MW satellites
      - Radio galaxies
      - Flat-spectrum radio quasars
      - ‘BL Lac’ objects
      - Gamma-ray Bursts



**Cosmic  
Particle  
Accelerators**

† 0.05-50 TeV

# Imaging Atm. Cherenkov Technique



## Atm. Cherenkov showers:

- V. large light pool ~250 m diameter
- Rapid time structure ~ 5 ns
- Very calorimetric
- Fine angular structure ( $< 1'$ )

## Imaging technique:

- Excellent shower reconstruction
- Large background rejection

Well-demonstrated by current instruments:  
H.E.S.S., MAGIC, & VERITAS

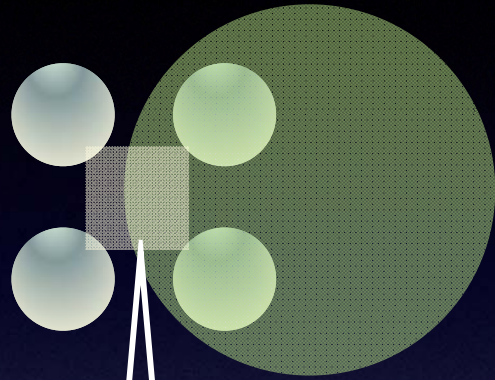
**But we have not reached limit of the technique !**

## Further improved by:

- More views of shower
- Higher resolution images
- Wider field-of-view

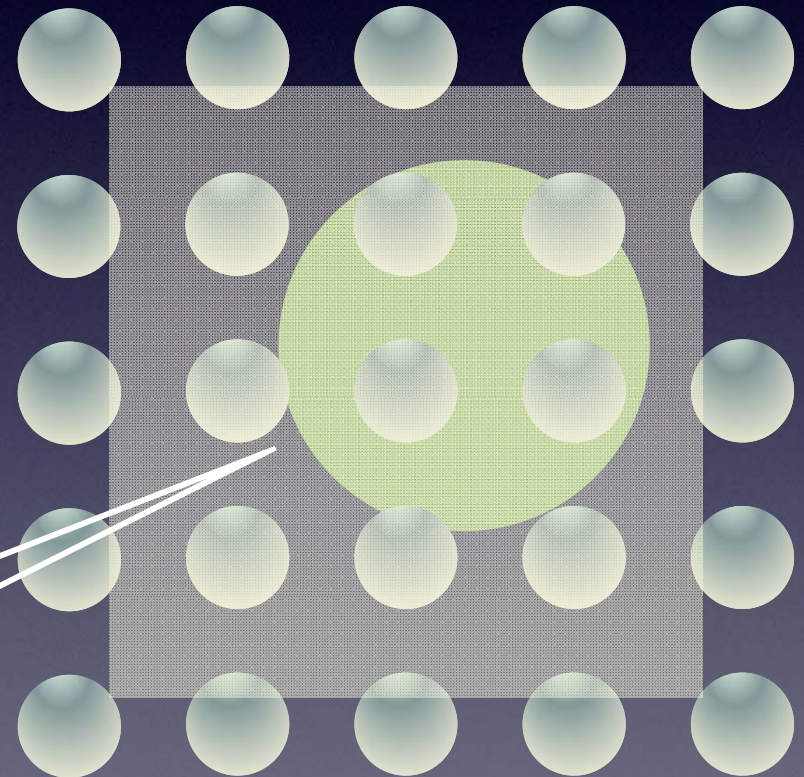


# Larger area → More contained events, more images



*Light pool radius*  
 $R \approx 100-150m$   
 $\approx$  typical telescope Spacing

*Sweet spot for best triggering & reconstruction...  
most showers miss it!*



- ✓ Larger detection Area
- ✓ More Images per shower
- ✓ Better  $\gamma$ -ray reconstruction
- ✓ Lower energy threshold

**What do we know, based on current instruments?**

**Great scientific potential exists in the VHE domain**

- *Frontier astrophysics & important connections to particle physics*

**Imaging Cherenkov technique is very powerful**

- *Have not yet reached its full potential → large telescope array*

**Exciting science in both Hemispheres**

- *Argues for an array in both S and N*

**Open Observatory gives substantial reward**

- *Open data/access, MWL connections to get the best science*

**International partnerships required by scale/scope**

- *Challenges associated with putting pieces together (i.e. funding streams, communities, etc.)*





# cta

cherenkov telescope array

# CTA Consortium



**The Consortium originated CTA and will contribute to the construction of the arrays**



**32 countries, ~1402 scientists, ~208 institutes, ~480 FTE**

# CTA Main Scientific Themes

## Cosmic Particle Acceleration

- How and where are particles accelerated?
- How do they propagate?
- What is their impact on the environment?



## Probing Extreme Environments

- Processes close to neutron stars and black holes
- Processes in relativistic jets, winds and explosions
- Exploring cosmic voids



## Physics frontiers – beyond the Standard Model

- What is the nature of Dark Matter? How is it distributed?
- Is the speed of light a constant for high-energy photons?
- Do axion-like particles exist?



→ See upcoming “Science with CTA” document





# Requirements & Drivers

**Energy coverage  
down to 20 GeV**  
*(Discovery domain:  
GRBs, Dark Matter)*

**Energy coverage  
up to 300 TeV**  
*(Pevatrons, hadron  
acceleration)*

**Good energy  
resolution, ~10-15%:**  
*(Lines, cutoffs)*

**Large Field of view 8-10°**  
*(Surveys, extended  
sources, flares)*

**Rapid Slew (20 s)  
to catch flares:**  
*(Transients)*

**10x Sensitivity &  
Collection Area**  
*(Nearly every topic)*

**Angular resolution < 0.1°  
above most of E range**  
*(Source morphology)*

# CTA Design (S array)

## Science Optimization under budget constraints

### Low energies

Energy threshold 20-30 GeV

23 m diameter

4 telescopes

**(LST's)**

### Medium energies

100 GeV – 10 TeV

9.7 to 12 m diameter

25 telescopes

**(MST's/SCTs)**

### High energies

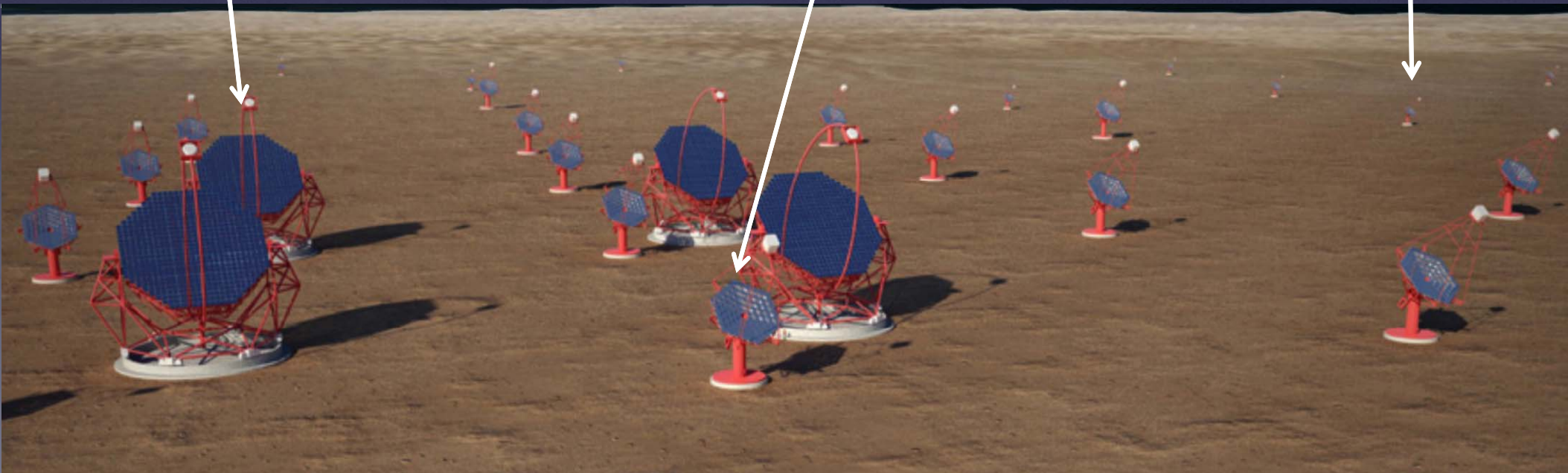
Up to > 300 TeV

10 km<sup>2</sup> eff. area @ 10 TeV

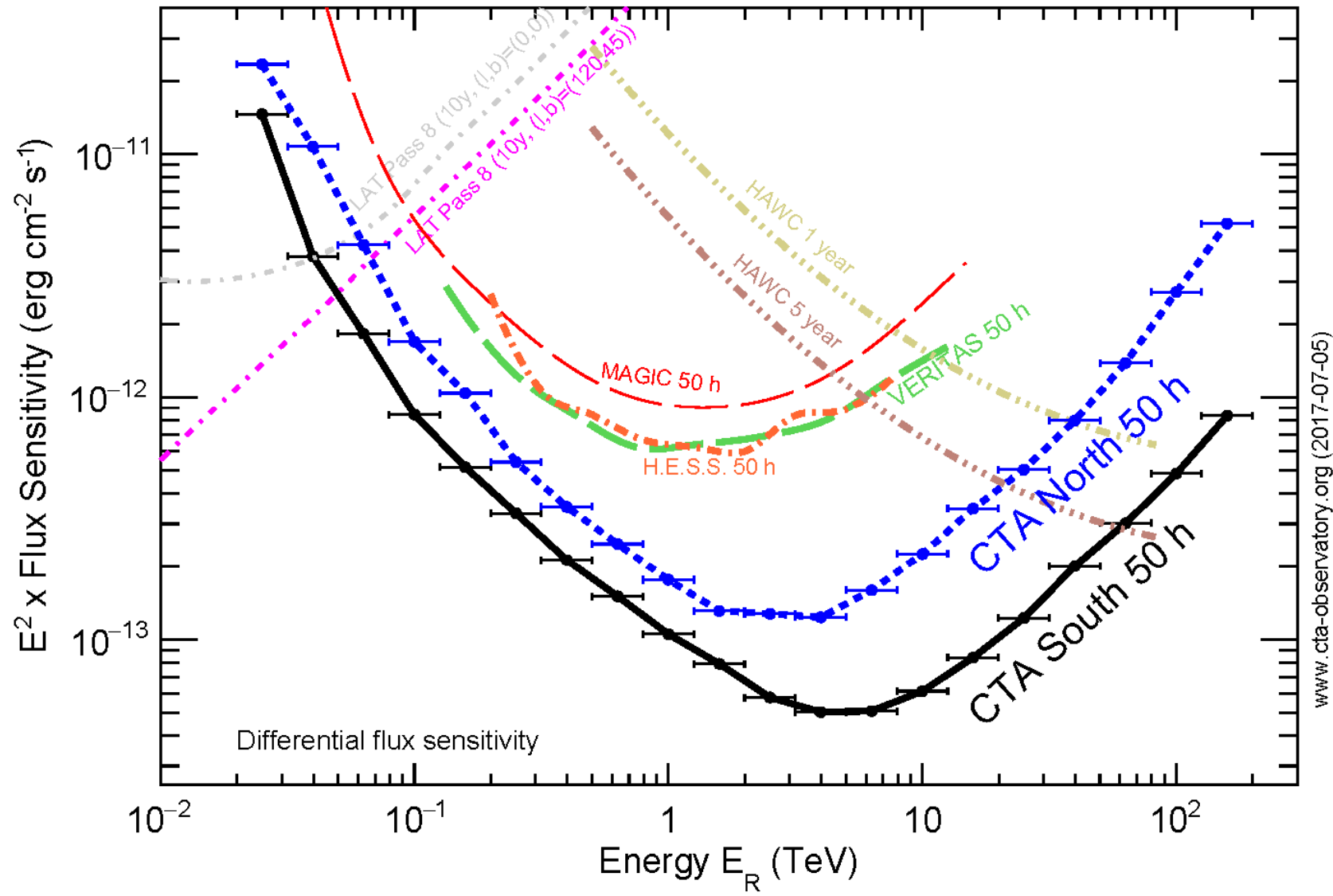
4m diameter

70 telescopes

**(SST's)**



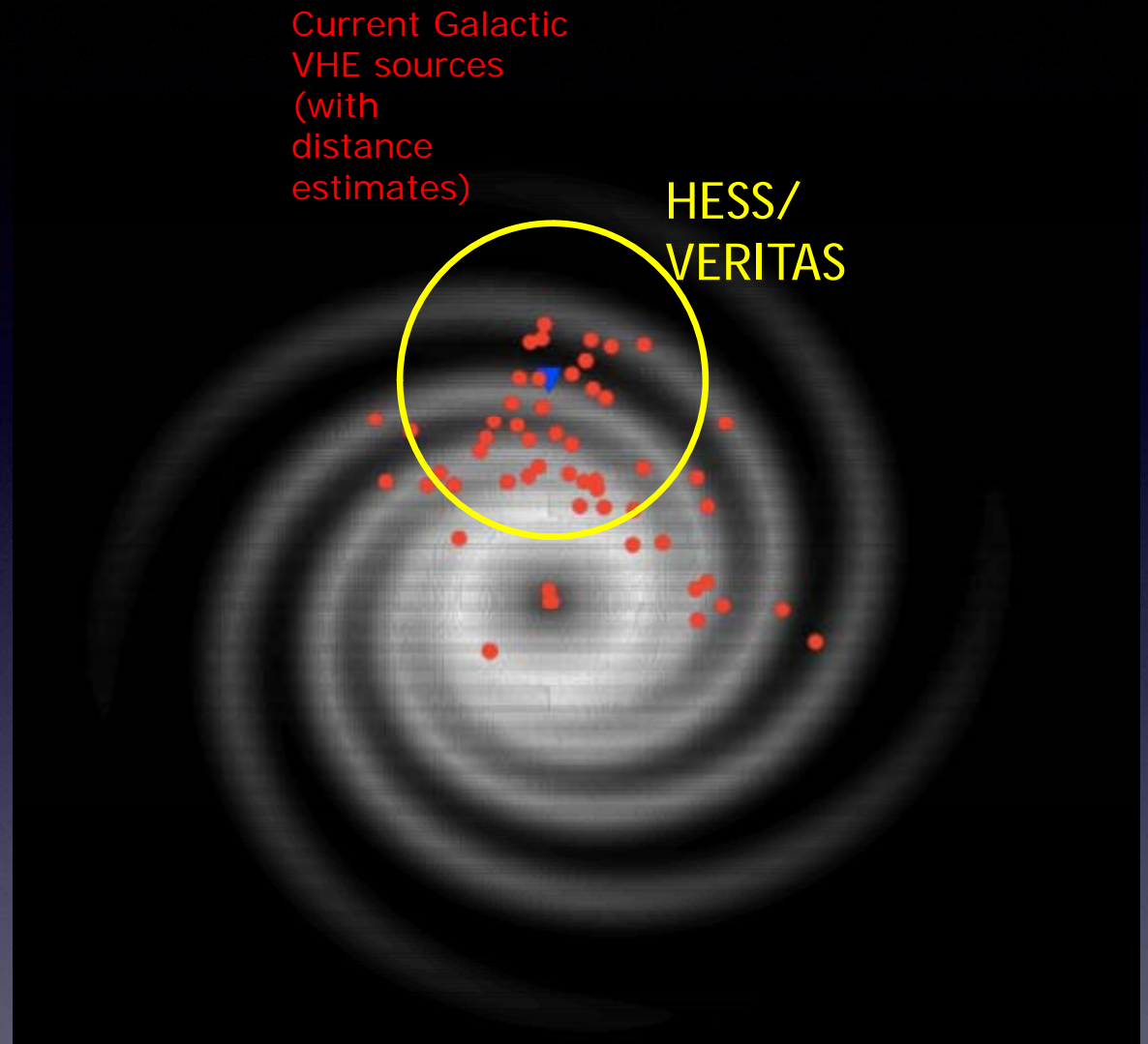
# Flux Sensitivity



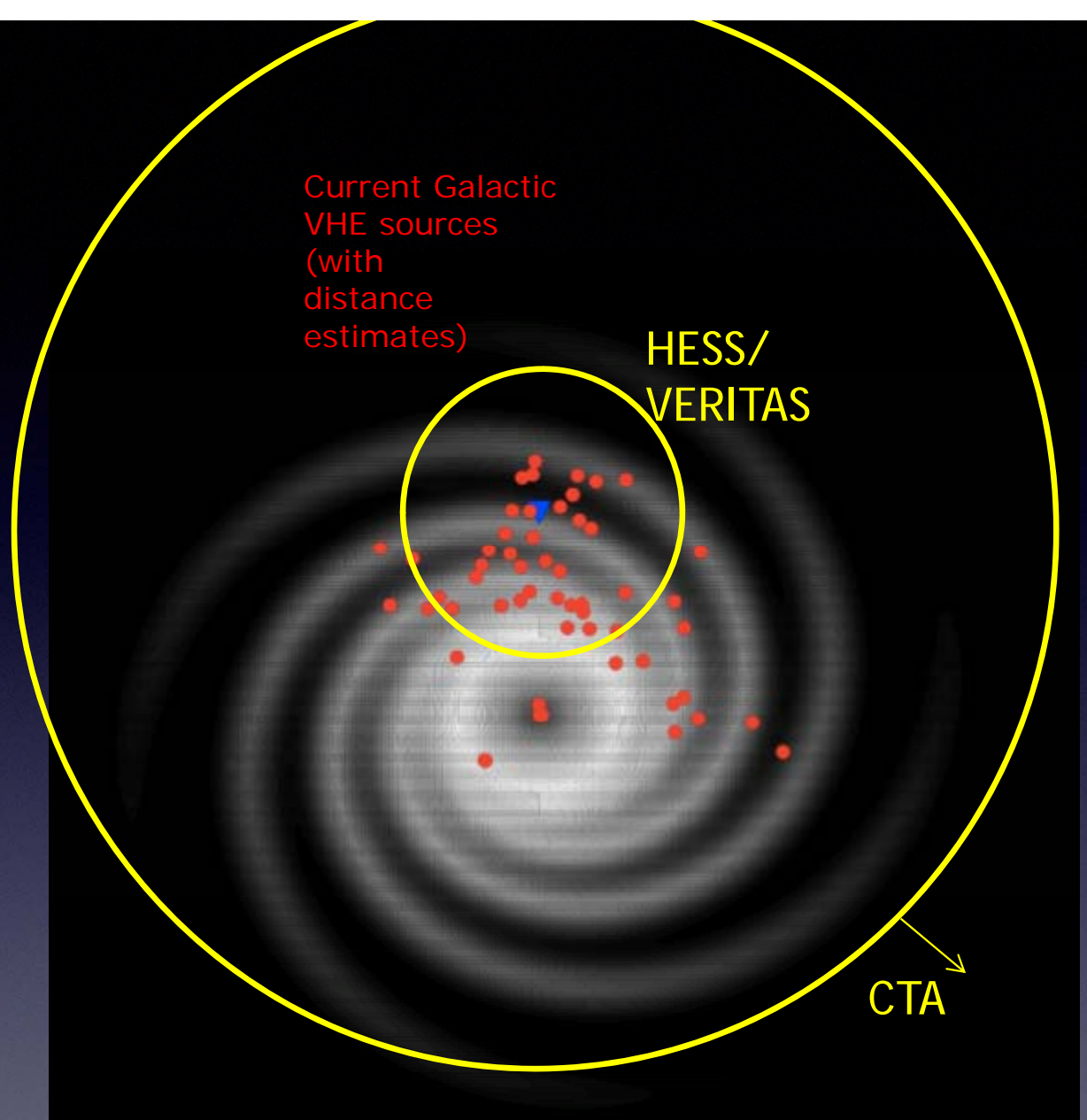
**Major sensitivity improvement & wider energy range**



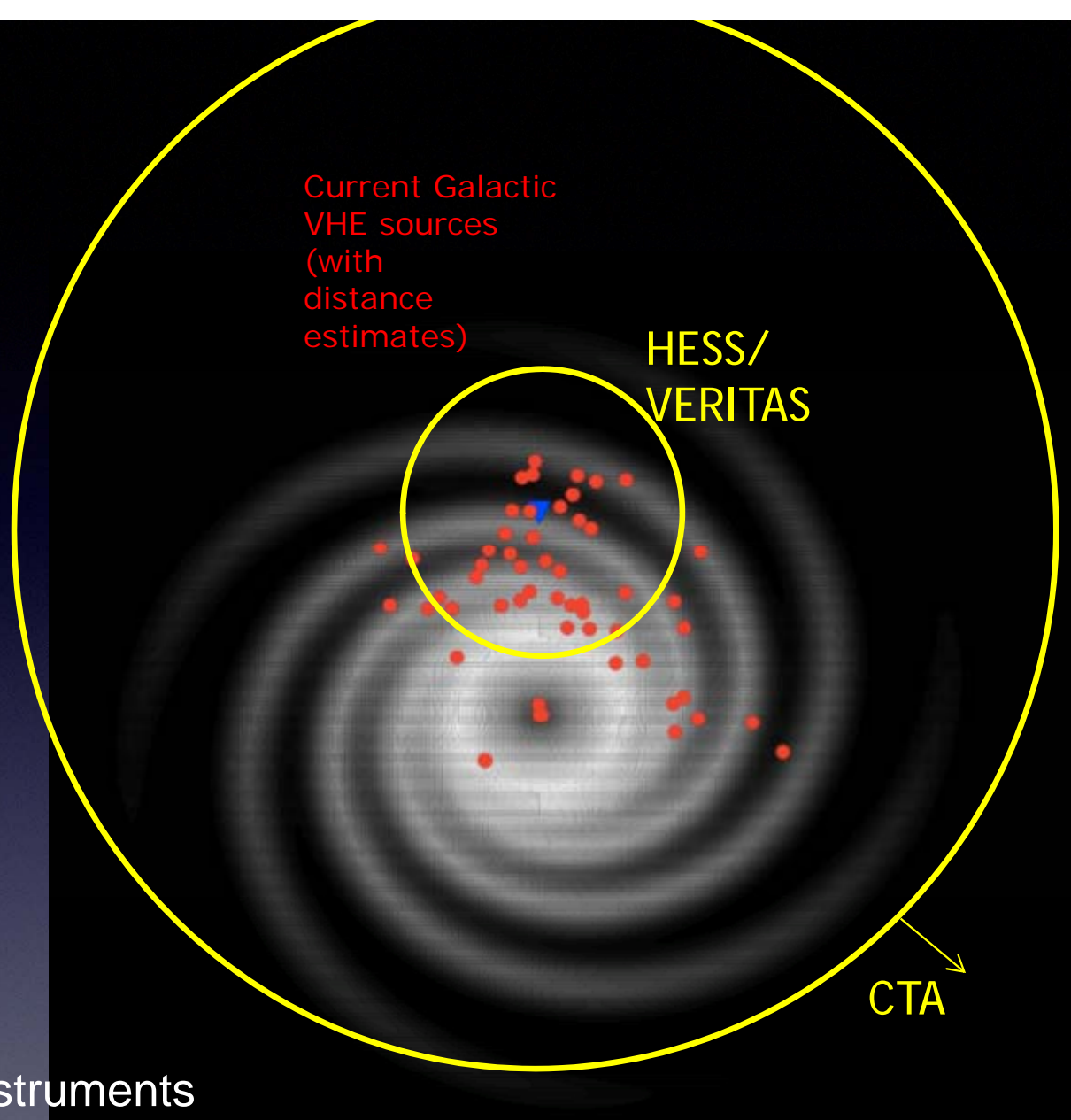
# Galactic Discovery Reach



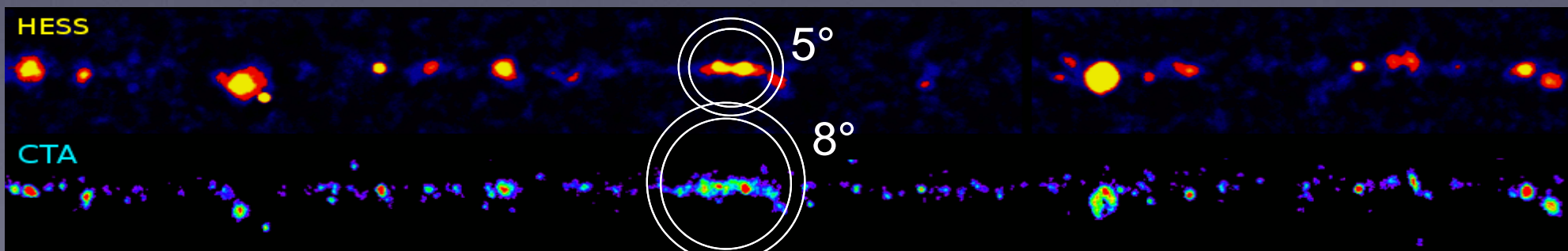
# Galactic Discovery Reach



# Galactic Discovery Reach



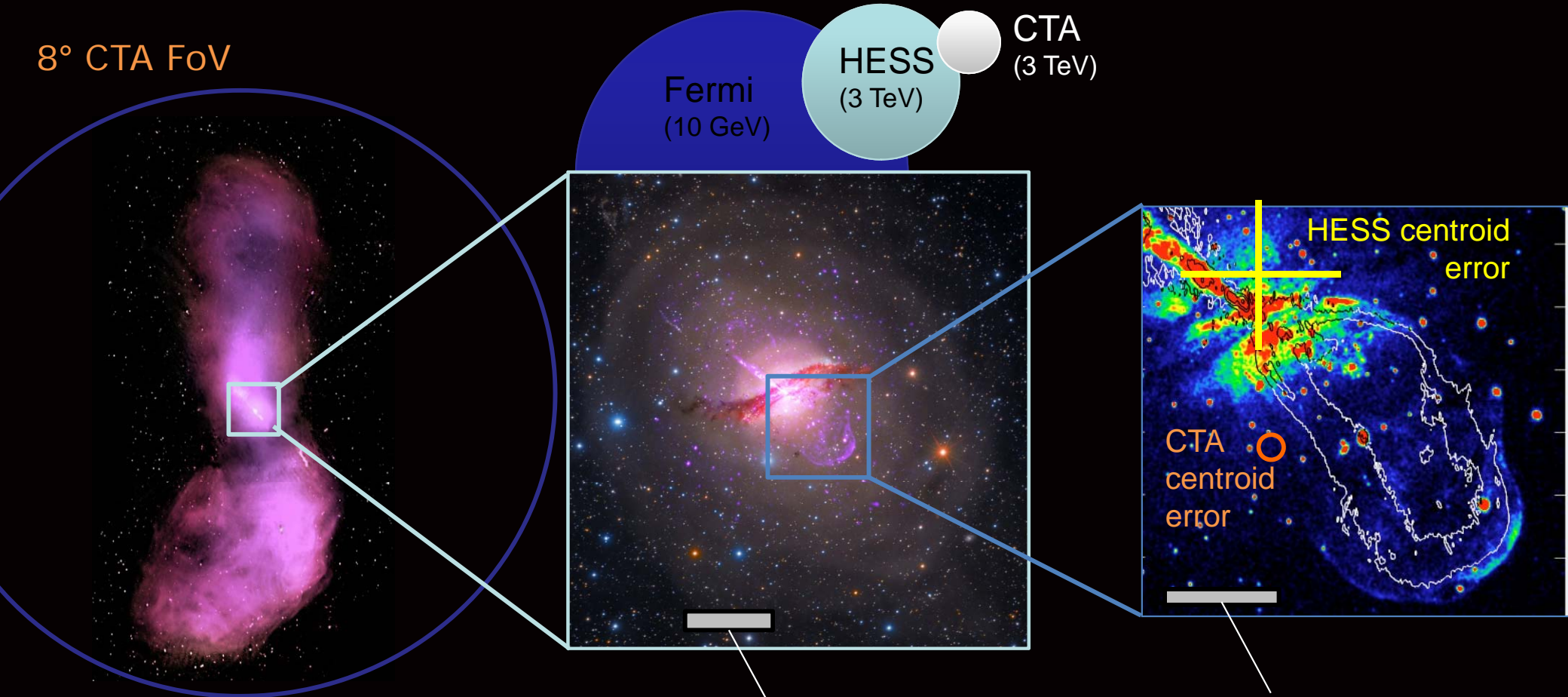
Survey speed:  
x300 faster than current instruments





# Angular Resolution

8° CTA FoV

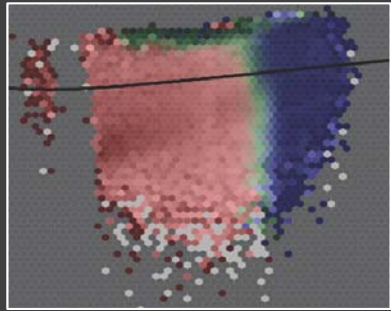


Example: Cen A

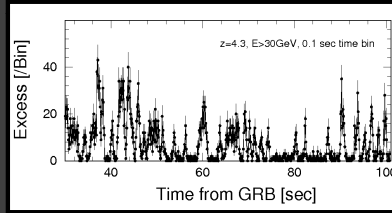
0.1°  
Typical  
HESS/MAGIC/VERITAS  
Resolution

2'  
CTA > 1 TeV

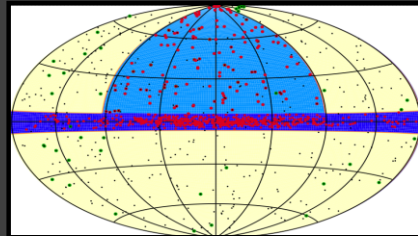
# Key Science Projects (KSPs)



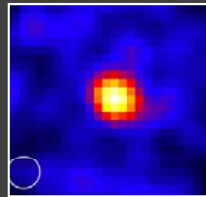
Dark Matter Programme



Transients



ExGal Survey

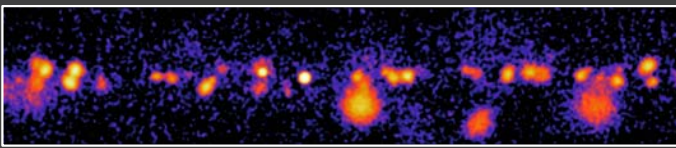
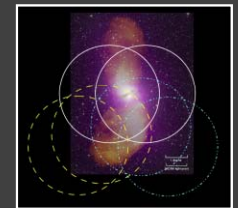


Galaxy Clusters



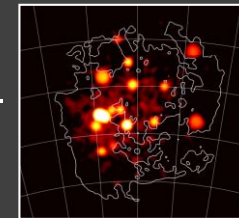
Star Forming Systems

AGN



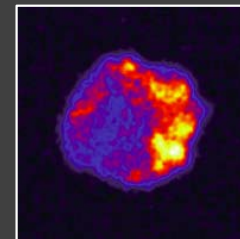
Galactic Plane Survey

LMC Survey

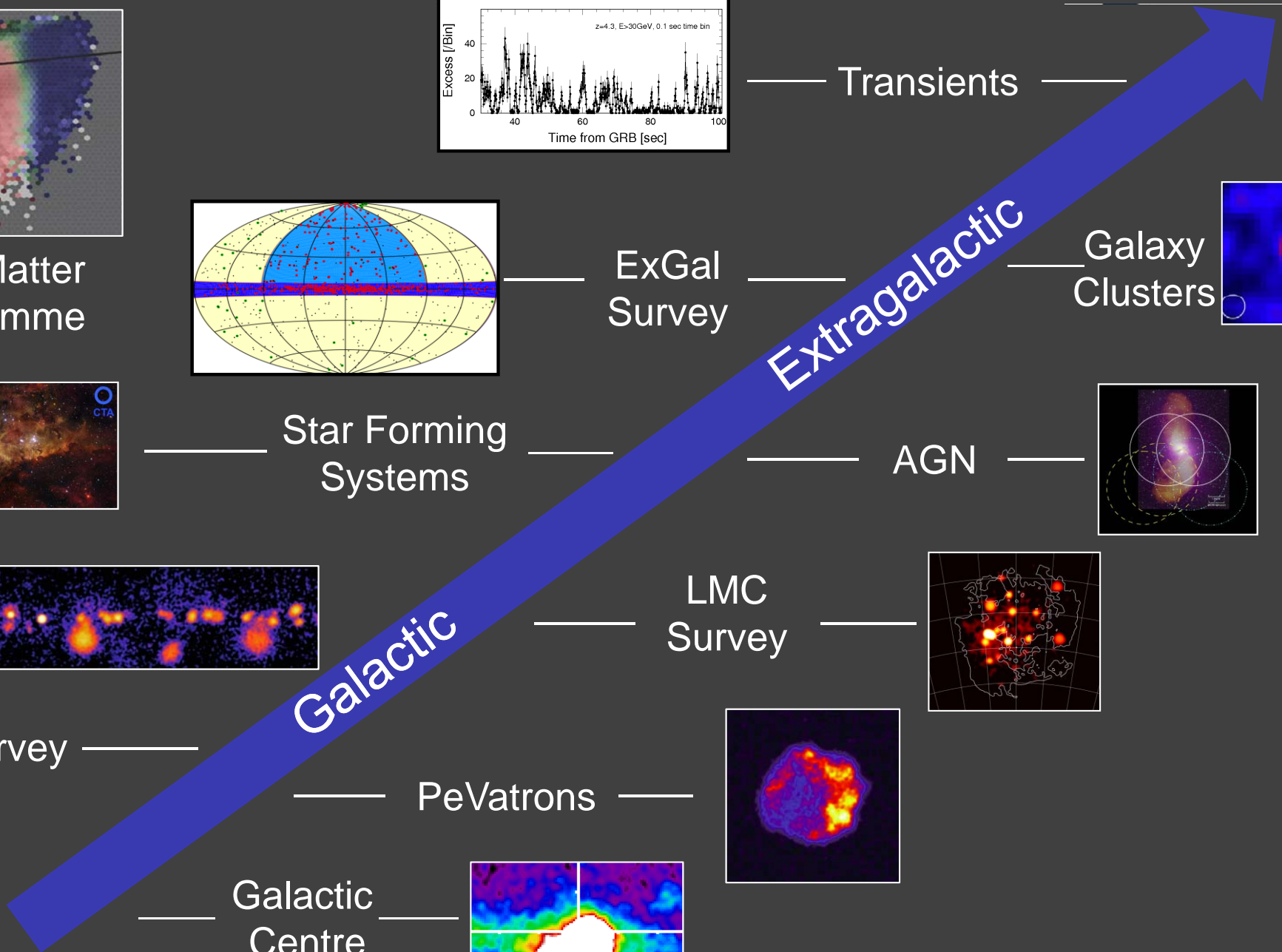
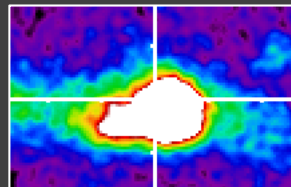


Galactic

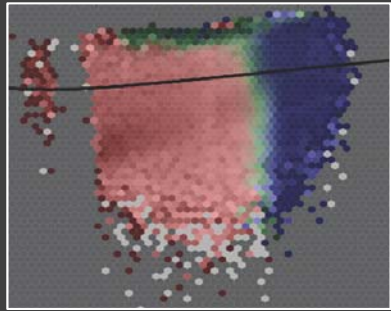
PeVatrons



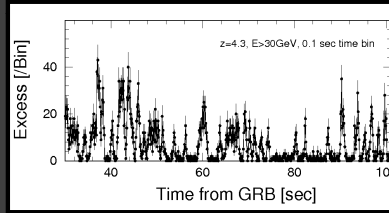
Galactic Centre



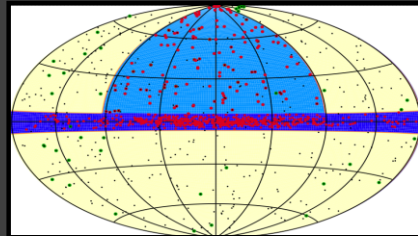
# Key Science Projects (KSPs)



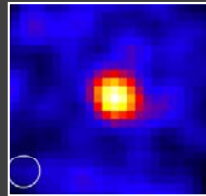
Dark Matter Programme



Transients



ExGal Survey

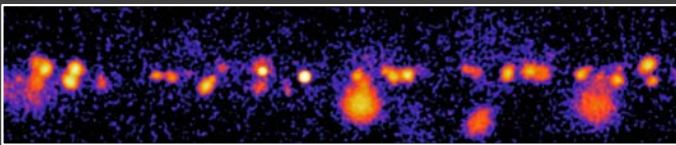
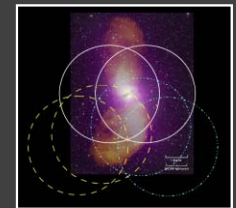


Galaxy Clusters



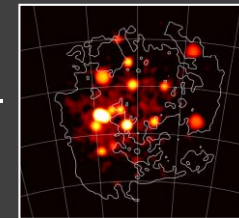
Star Forming Systems

AGN



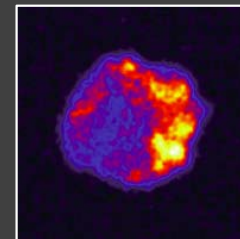
Galactic Plane Survey

LMC Survey

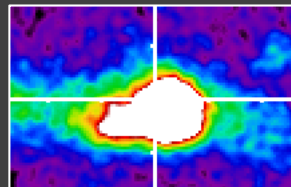


Galactic

PeVatrons



Galactic Centre



Extragalactic

CTA Science talks:  
R. Zanin 15/07 GA044  
T. Hassan 18/07 GA145  
A. Morselli 19/07 DM015



# Large Telescope (LST)

23 m diameter  
390 m<sup>2</sup> dish area  
28 m focal length  
1.5 m mirror facets

4.5° field of view  
0.1° PMT pixels  
Camera  $\emptyset$  over 2 m

*Carbon-fiber structure  
for 20 s positioning*

Active mirror control

**4 LSTs on South site  
4 LSTs on North site**

**Prototype construction  
Underway (La Palma)**

Talk by M. Teshima – this session



# Medium Telescope (MST)

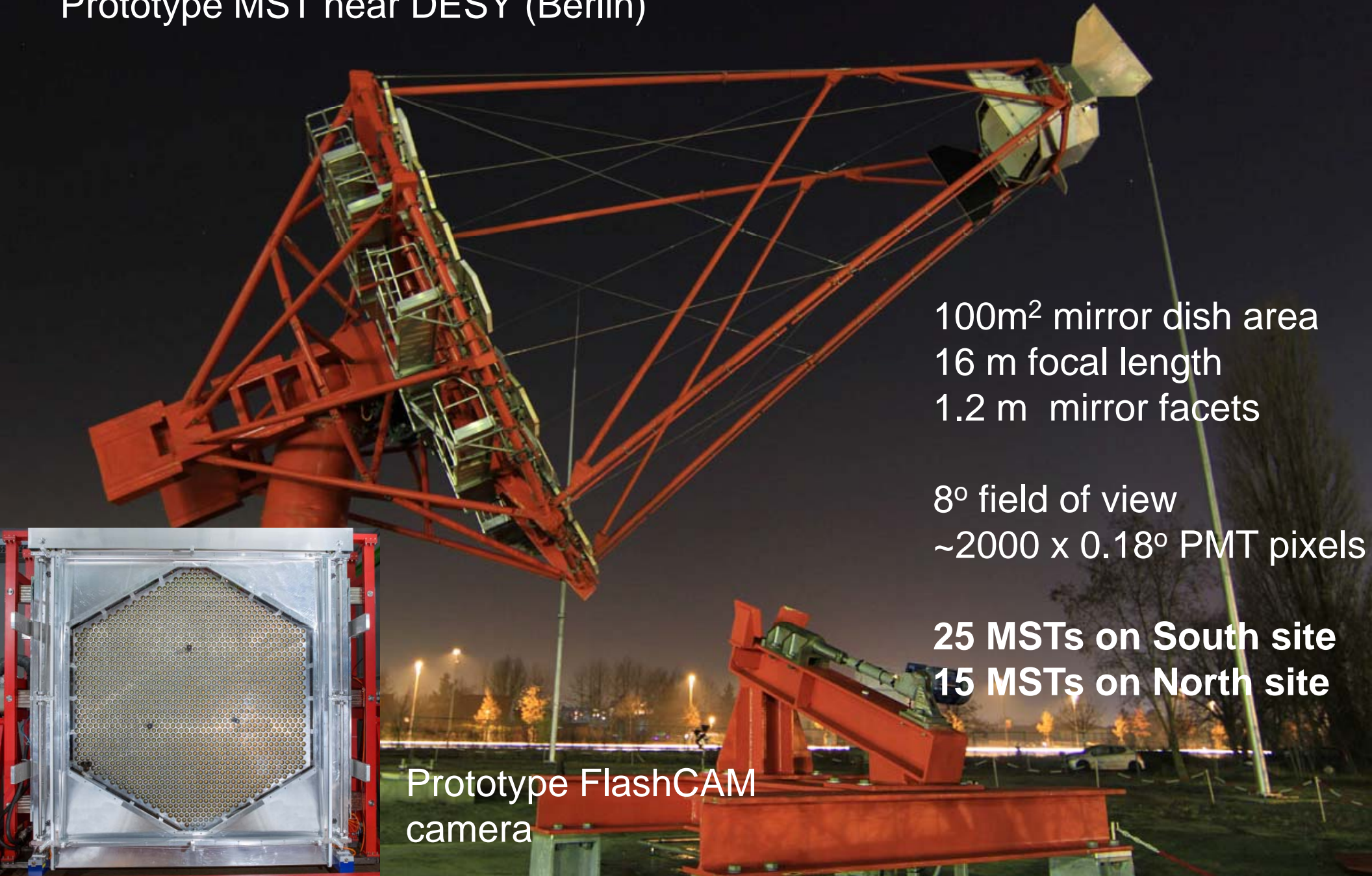
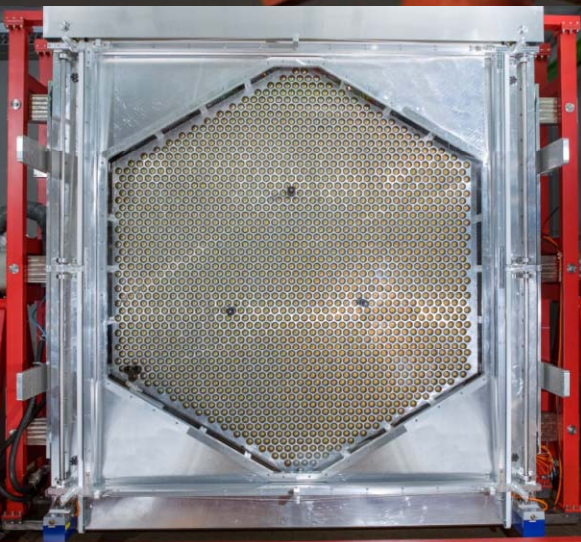
Prototype MST near DESY (Berlin)

100m<sup>2</sup> mirror dish area  
16 m focal length  
1.2 m mirror facets

8° field of view  
~2000 x 0.18° PMT pixels

25 MSTs on South site  
15 MSTs on North site

Prototype FlashCAM  
camera



# Medium 2-mirror Telescope



Prototype SCT at Whipple Obs, Arizona

Schwarzschild-Couder Telescope (SCT)

9.7 m primary

5.4 m secondary

5.6 m focal length,  $f/0.58$

50 m<sup>2</sup> mirror dish area

PSF better than 4.5'  
across 8° FOV

8° field of view

11328 x 0.07° Si-PM pixels

→ Improved  $\gamma$ -ray angular resolution

Talk by V. Vassiliev  
– this session



# Small Sized Telescopes (SSTs)

- 3 different prototype designs
- 2 designs use two-mirror approaches (Schwarzschild-Couder design)
- All use Si-PM photosensors
- 8-10 m<sup>2</sup> mirror area, FOV > 9°



SST-1M  
Krakow, Poland

Talk by C. Alispach  
– this session



SST-2M ASTRI  
Mt. Etna, Italy

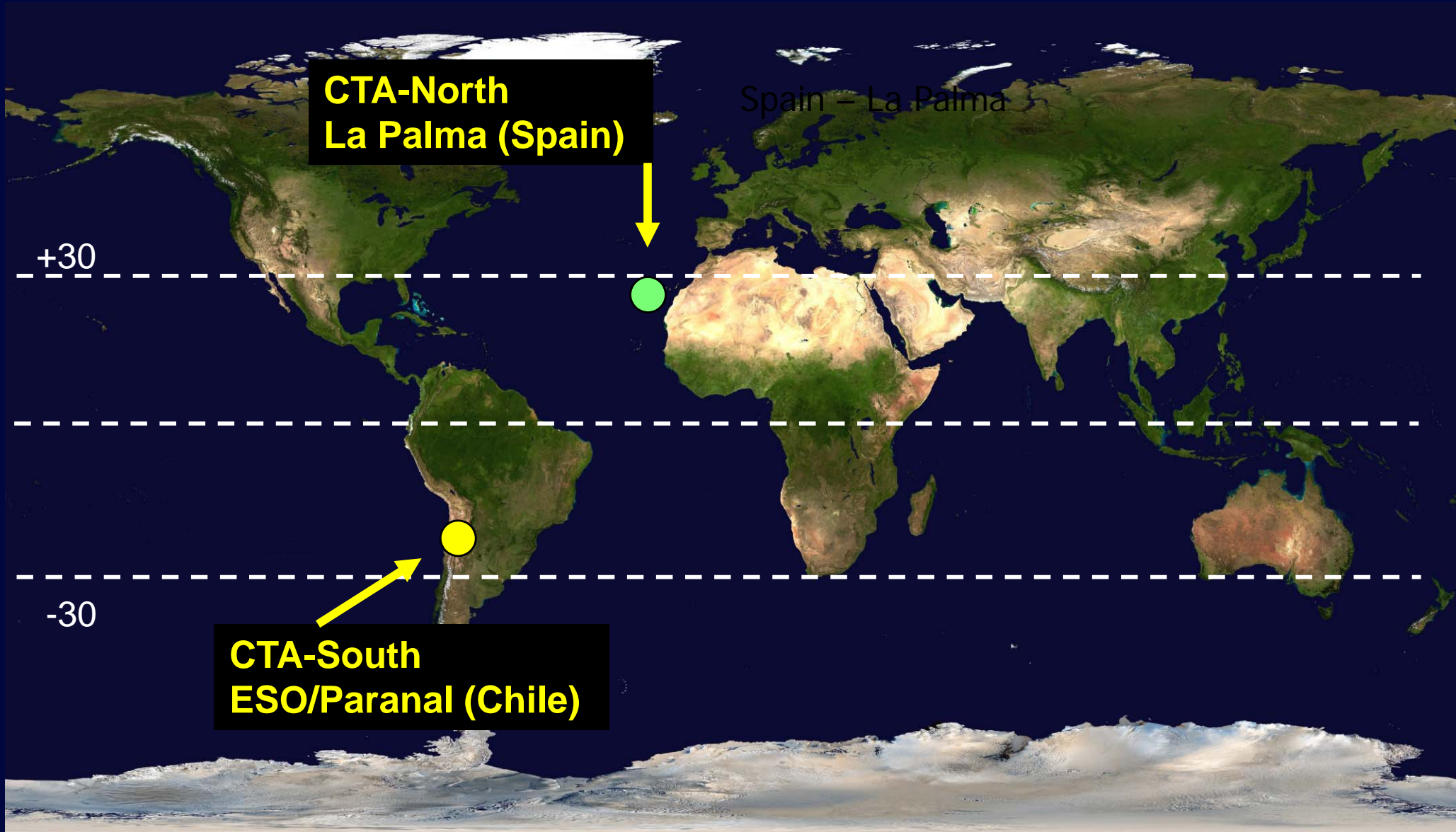
Talk by M.C. Maccarone  
– this session



SST-2M GCT  
Meudon, France

Talk by H. Sol  
– Monday, 13:30-15:00

# CTA Sites

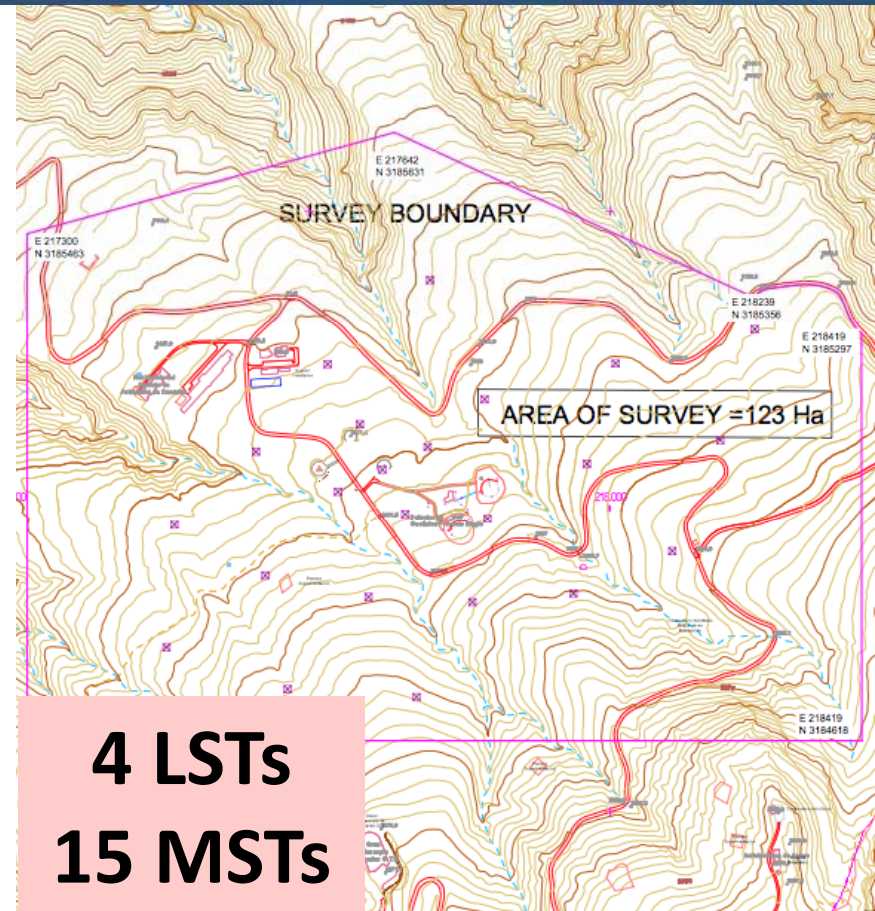




# La Palma – CTA North



- Canary Islands, Spain
- Observatorio del Roque de los Muchachos
- Existing observatory, under management by Instituto de Astrofísica de Canarias (IAC)
- Site of LST 1 & existing MAGIC telescopes
- Current work: topographical study, building concepts, tender for geotechnical study soon





# ESO PARANAL – CTA South

- Atacama Desert, Chile, south of Cerro Paranal
- Existing observatory, under management by European Southern Observatory (ESO)
- Near a set of existing (VLT) and future (ELT) telescopes

Vulcano Lullillaco  
6739 m, 190 km east

Cerro Armazones  
E-ELT

Cherenkov Telescope Array Site

Cerro Paranal  
Very Large Telescope

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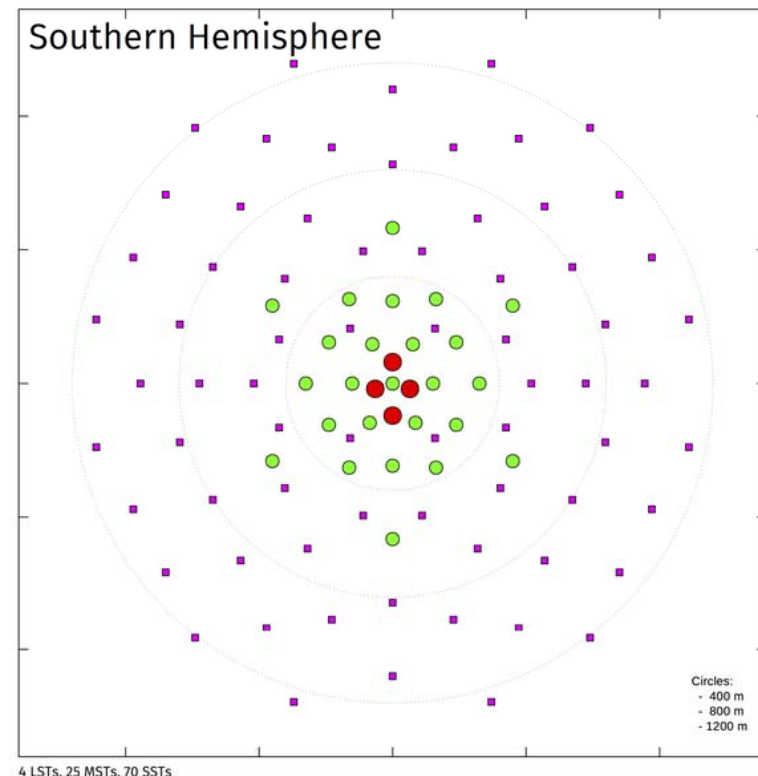
Vulcano Lullillaco  
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E-ELT

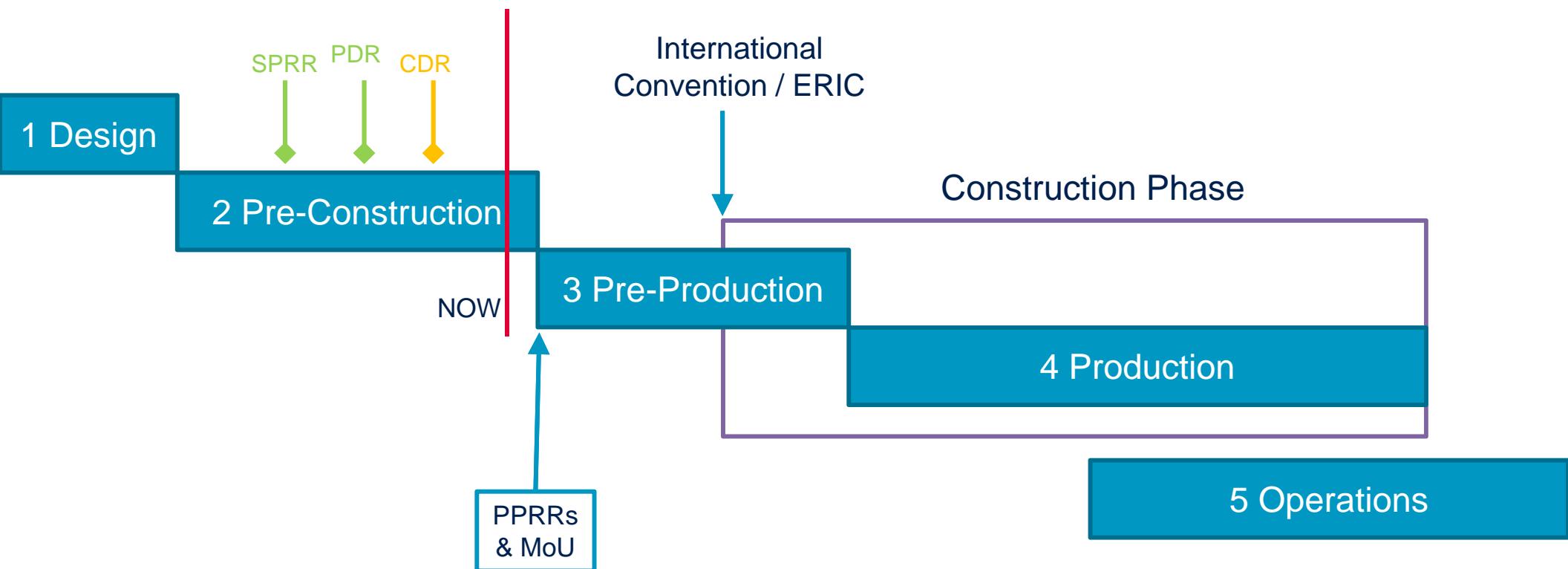
Current work: geotechnical studies (boreholes), topographical survey, concepts for roads, power, ducting, & buildings

Cerro Paranal  
Very Large Telescope

**4 LSTs**  
**25 MSTs**  
**70 SSTs**



# CTA Phases & Timeline



- 2016-7: Hosting agreements, site preparations start
- 2018: Start of construction
- Funding level at ~65% of required for *baseline implementation*
  - start with *threshold implementation*
  - additional funding & telescopes needed to complete baseline CTA
- Construction period of ~6 years
- Initial science with partial arrays possible before construction end



- **We've learned a lot from previous/present experiments**

With many discoveries, VHE  $\gamma$ -ray astronomy has become a major and exciting field of research

Outstanding science potential and the power of the atmospheric Cherenkov technique  $\rightarrow$  CTA

- **Cherenkov Telescope Array (CTA)\***

*Outstanding sensitivity & resolution over wide energy range*

*Far-reaching key science program*

*Open observatory with all data released to public*

- CTA prototyping/design is largely completed; now ready to develop both sites and enter pre-production of telescopes

- **In next decade, CTA will provide data of a quality not yet seen in the HE/VHE  $\gamma$ -ray band**

**\*We gratefully acknowledge financial support from the agencies and organizations listed here: [http://www.cta-observatory.org/consortium\\_acknowledgments](http://www.cta-observatory.org/consortium_acknowledgments).**

# CTA Talks at ICRC 2017



## GA Parallel, Friday, July 14: 16:30-18:30

|                          |       |  |
|--------------------------|-------|--|
| Rene Ong                 | GA325 | Cherenkov Telescope Array: The Next Generation Gamma-ray Observatory   |
| Masahiro Teshima         | GA202 | Large Size Telescope of the Cherenkov Telescope Array  |
| Cyril Alispach           | GA300 | Performance of a small size telescope (SST-1M) camera for gamma-ray astronomy with the Cherenkov Telescope Array |
| Maria Concetta Maccarone | GA022 | ASTRI for the Cherenkov Telescope Array  |
| Vladimir Vassiliev       | GA051 | Prototype 9.7m Schwarzschild-Couder telescope for the Cherenkov Telescope Array: Project Overview                |
| Jan Ebr                  | GA077 | Atmospheric calibration of the Cherenkov Telescope Array   |

## GA Parallel, Saturday, July 15: 13:30-15:00

|               |       |   |
|---------------|-------|---|
| Roberta Zanin | GA044 | Observing the Galactic Plane with the Cherenkov Telescope Array |
|---------------|-------|---|

## GA Parallel, Monday, July 17: 13:30-15:00

|            |       |  |
|------------|-------|--|
| Helene Sol | GA123 | Observing the sky at extremely high energies with CTA: Status of the GCT project |
|------------|-------|--|

## GA Parallel, Tuesday, July 18: 16:30-18:30

|              |       |  |
|--------------|-------|--|
| Tarek Hassan | GA145 | Extragalactic source population studies at very high energies in the Cherenkov Telescope Array era |
|--------------|-------|--|

## GA Parallel, Wednesday, July 19: 16:30-18:30

|             |       |   |
|-------------|-------|---|
| David Kieda | GA094 | Stellar Intensity Interferometric Capabilities of IACT Arrays |
|-------------|-------|---|

## DM Parallel, Wednesday, July 19: 16:30-18:30

|               |       |  |
|---------------|-------|--|
| Aldo Morselli | DM015 | The Dark Matter Programme of the Cherenkov Telescope Array |
|---------------|-------|--|

# CTA Posters at ICRC 2017 I



[CRI097] A Monte Carlo simulation study for cosmic-ray chemical composition measurement with Cherenkov Telescope Array

Board #: 147

Presented by Michiko OHISHI on 18 Jul 2017 at 15:00

[GA019] Design, development and characterization of a calibration system for the camera of the Large Size Telescope proposed for CTA

Board #: 173

Presented by Michele PALATIELLO on 13 Jul 2017 at 15:00

[GA021] Atmospheric monitoring and array calibration in CTA using the Cherenkov Transparency Coefficient

Board #: 179

Presented by Stanislav STEFANIK on 13 Jul 2017 at 15:00

[GA023] Tools and Procedures for the CTA Array Calibration

Board #: 185

Presented by Maria Concetta MACCARONE on 13 Jul 2017 at 15:00

[GA024] Sun/Moon photometer for Cherenkov Telescope Array \u2013 first results

Board #: 240

Presented by Jakub JURYSEK on 13 Jul 2017 at 15:00

[GA039] Performance of the Cherenkov Telescope Array

Board #: 150

Presented by Gernot MAIER on 13 Jul 2017 at 15:00

[GA040] Raman LIDARs for atmospheric calibration in CTA

Board #: 230

Presented by Georges VASILEIADIS on 13 Jul 2017 at 15:00

[GA041] Control Software for a Small-Size Telescope (SST-1M) proposed for the Cherenkov Telescope Array

Board #: 188

Presented by Roland WALTER on 13 Jul 2017 at 15:00

[GA042] End-to-end data acquisition pipeline for the Cherenkov Telescope

Board #: 187

Presented by Roland WALTER on 13 Jul 2017 at 15:00



# CTA Posters at ICRC 2017 II



[GA059] Studies of the nature of the low-energy, gamma-like background for Cherenkov Telescope Array

Board #: 214

Presented by Julian SITAREK on 13 Jul 2017 at 15:00

[GA061] Towards final characterization and performance of the GCT prototype telescope structure for CTA

Board #: 176

Presented by Cedric PERENNES on 13 Jul 2017 at 15:00

[GA102] Searching for PeVatrons in the CTA Galactic Plane Survey

Board #: 149

Presented by Cyril TRICHARD on 13 Jul 2017 at 15:00

[GA131] A Compact High Energy Camera (CHEC) for the GCT of CTA

Board #: 183

Presented by Harm SCHOORLEMMER on 13 Jul 2017 at 15:00

[GA136] Prototype 9.7m Schwarzschild-Couder telescope for the Cherenkov Telescope Array: status of the optical system

Board #: 209

Presented by Daniel NIETO on 13 Jul 2017 at 15:00

[GA141] Baseline telescope layouts of the Cherenkov Telescope Array

Board #: 233

Presented by Paolo CUMANI on 13 Jul 2017 at 15:00

[GA146] Exploring deep learning as an event classification method for the Cherenkov Telescope Array

Board #: 210

Presented by Daniel NIETO on 13 Jul 2017 at 15:00

[GA147] A Trigger Interface Board to manage trigger and timing signals in CTA Large-Sized Telescope and Medium-Sized Telescope camera

Board #: 208

Presented by Marcos LOPEZ on 13 Jul 2017 at 15:00

[GA155] ASTRI SST-2M prototype and mini-array simulation chain, data reduction software, and archive in the framework of the CTA

Board #: 184

Presented by Maria Concetta MACCARONE on 13 Jul 2017 at 15:00

# CTA Posters at ICRC 2017 III



[GA158] A pointing solution for the medium size telescopes for the Cherenkov Telescope Array

Board #: 186

Presented by Domenico TIZIANI on 13 Jul 2017 at 15:00

[GA185] Studying cosmological gamma-ray propagation with the Cherenkov Telescope Array

Board #: 099

Presented by Florian GATÉ on 13 Jul 2017 at 15:00

[GA279] Gammapy - high level data analysis for extragalactic science cases with the Cherenkov Telescope Array

Board #: 118

Presented by Julien LEFAUCHEUR on 13 Jul 2017 at 15:00

[GA284] The ARCADE Raman Lidar and atmospheric simulations for the Cherenkov Telescope Array

Board #: 231

Presented by Laura VALORE on 13 Jul 2017 at 15:00

[GA166] Development of a strategy for calibrating the novel SiPM camera of the SST-1M telescope proposed for the Cherenkov Telescope Array

Board #: 238

Presented by Imen AL SAMARAI on 13 Jul 2017 at 15:00

[GA278] Gammapy - A prototype for the CTA science tools

Board #: 215


Presented by Matteo CERRUTI on 13 Jul 2017 at 15:00

# Visit the CTA Exhibit!



cherenkov telescope array

## A Ground-Based Gamma-Ray Astronomy Observatory



an observatory for ground-based gamma-ray astronomy

about 



CTA will be the largest and most sensitive ground-based gamma-ray observatory in the world with more than 100 telescopes located in the northern and southern hemispheres.


CTA's sensitivity to energies up to 300 TeV will push CTA beyond the edge of the known electromagnetic spectrum, providing a previously unexplored view and allowing us to search for extreme particle acceleration.

CTA's detection of gamma rays with energies as low as 20 GeV will allow CTA to probe transient and time-variable gamma-ray phenomena in the very distant Universe with unprecedented sensitivity.

CTA will be the first ground-based gamma-ray observatory open to the world community and a resource for data from unique, high-energy astronomical observations.




an observatory for ground-based gamma-ray astronomy

science 

CTA will seek to understand the impact of high-energy particles in the evolution of cosmic systems and to gain novel insight into the most extreme phenomena in the Universe.


### Exploring the most extreme phenomena in the Universe

CTA will build on the advances pioneered by its predecessors to expand the catalogue of known gamma-ray emitting cosmic sources tenfold, detecting more than 1,000 new objects.



CTA will seek to address questions hiding under three major themes: **uncovering the origin and role of Relativistic Cosmic Particles, Probing Extreme Environments and Pushing Frontiers in Physics.**

CTA will conduct a **search of particle acceleration** in the Universe and will search for annihilating dark matter particles and deviations from Einstein's theory of special relativity.



an observatory for ground-based gamma-ray astronomy

technology 



CTA's telescopes will detect gamma rays by capturing the Cherenkov light that is produced when they interact with the Earth's atmosphere. The mirrors reflect the light to the camera, which capture the event and convert it into an electrical signal that is digitized and transmitted to record the image of the light.

CTA's three classes of telescopes – the Large-Sized Telescope, Medium-Sized Telescope and Small-Sized Telescope – will provide broad energy coverage from billions to trillions times the energy of visible light (20 GeV to 300 TeV).

CTA's telescope structures will stand between about 10 and 45 metres tall and weigh between 8 and 100 tonnes.

CTA will be 10 times more sensitive than existing instruments and will look at the gamma-ray sky with a resolution and sensitivity that is unprecedented.



an observatory for ground-based gamma-ray astronomy

cherenkov telescope array

## Exploring the Universe at the Highest Energies



an observatory for ground-based gamma-ray astronomy

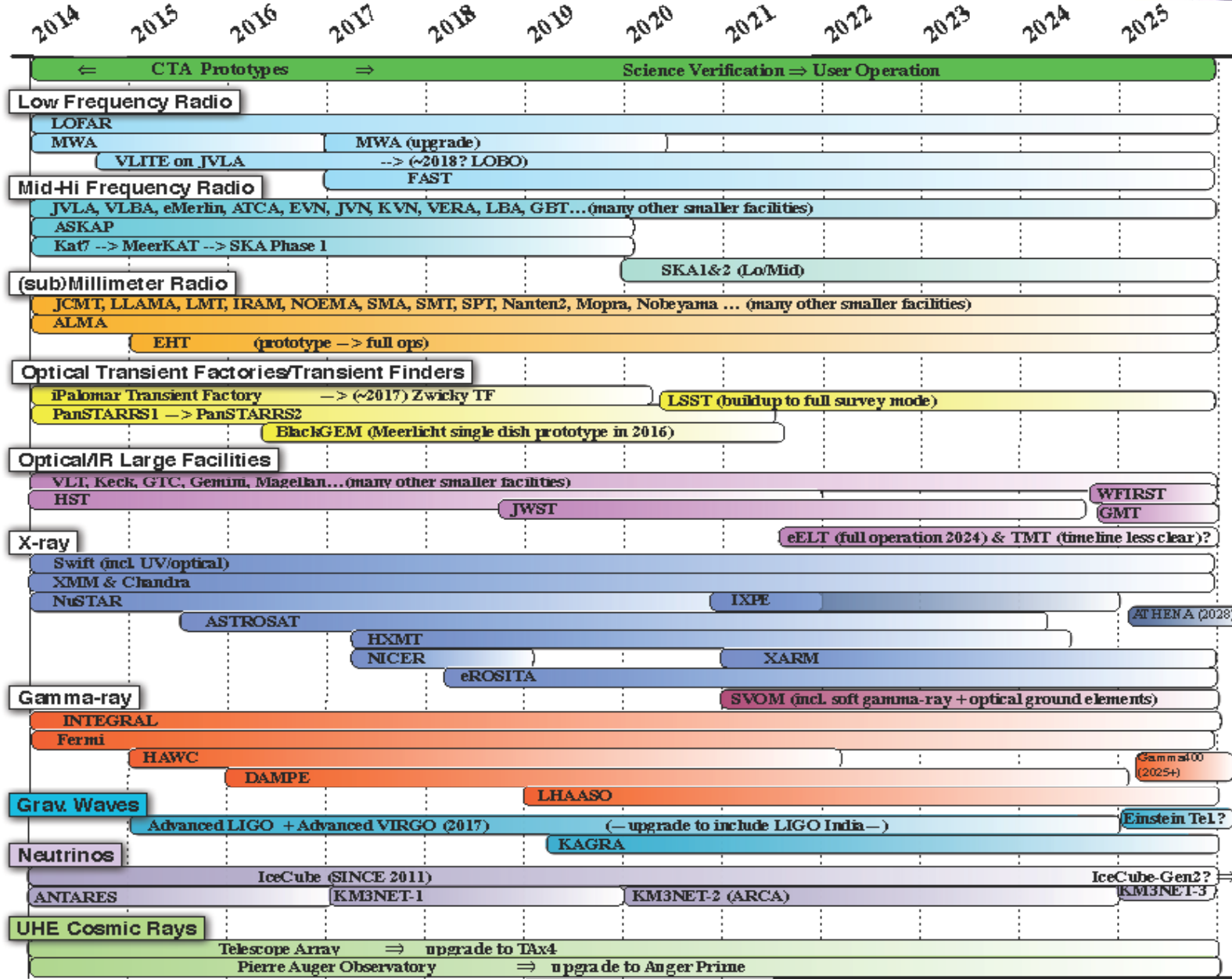


# BACKUP



cherekov  
telescope  
array

# Important MWL/MM Synergies



Caveat: Observatory timelines are very uncertain; this represents a notional picture based on available information



# Science with the Cherenkov Telescope Array

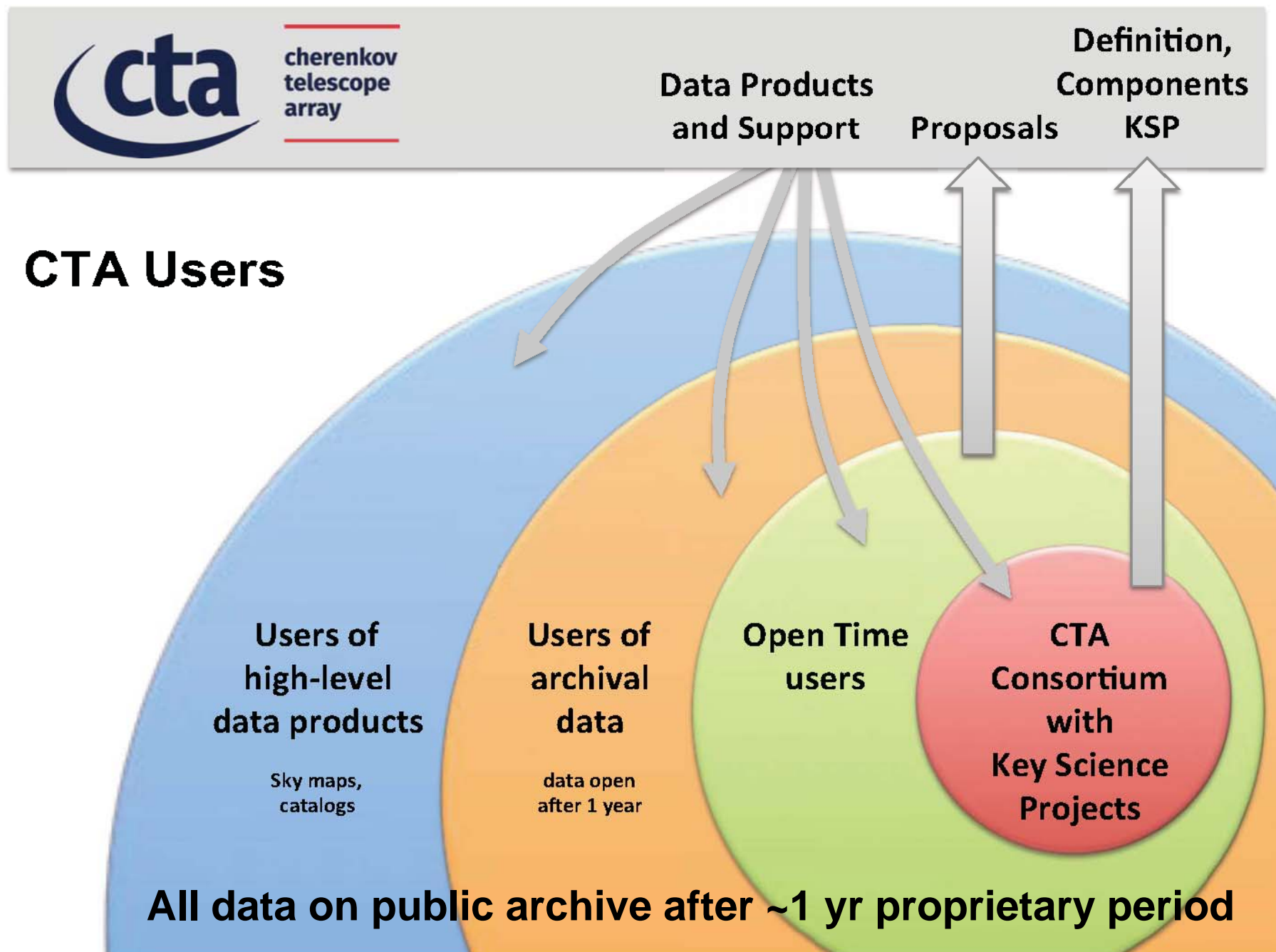
## Science with CTA

200 page document describing  
core CTA science

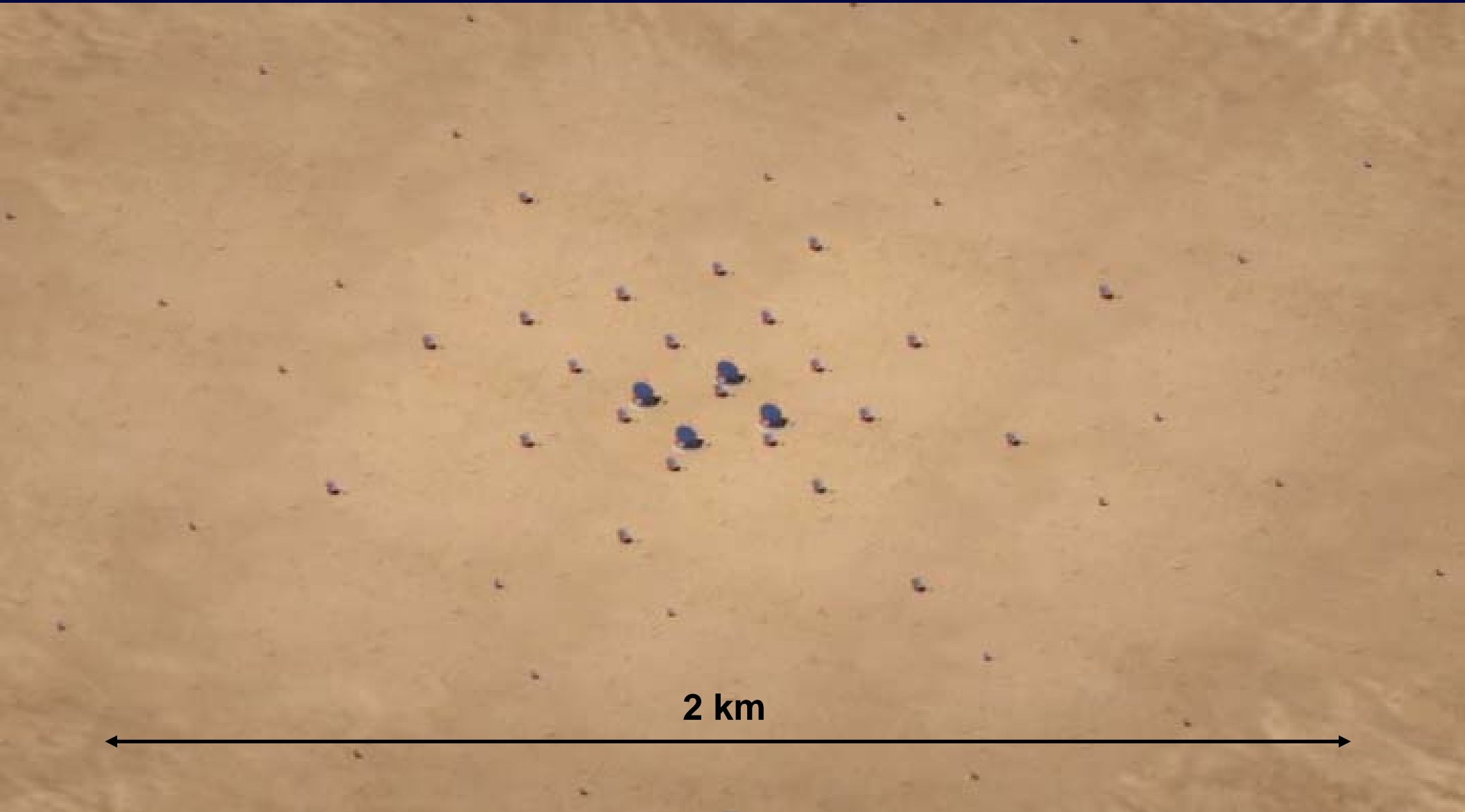
Will soon be put on arXiv and  
become a regular book



# CTA: An Open Observatory



# CTA South Array



2 km

# Telescope Types

